

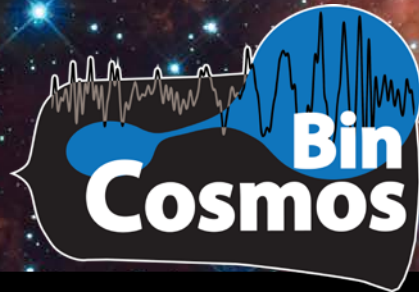


Background design E. Buunk. Credit: NASA/Paresce



\*Field, Non  
Common Envelope

Selma E. de Mink  
**Black-hole Binary Formation\***  
Sackler Conference 2018



With credit to the following external collaborators: **Chris Belczynski, Ilya Mandel, Matteo Cantiello, Andrew King, Norbert Langer, Pablo Marchant, Philip Podsiadlovski, Hugues Sana, Simon Stevenson, Ed van den Heuvel, Alejandro Vigna-Gómez, Nathan Smith, Maria Drout, Thomas Kupfer, Sung-Chul Yoon, Wolfgang Kerzendorf, Jose Groh, Carles Badenes, Simon Portegies-Zwart, Rob Izzard, Coen van Neijssel, Abel Schootemeijer, Onno Pols, Tom Maccarone, Chris Evans, Ori Fox, Schuyler van Dyk, Claus Leitherer, Leonardo Almeida, Alex de Koter, Tony Piro, Paul Crowther, Danny Lennon, Imants Platais, VLT-FLAMES Massive Star Consortium**

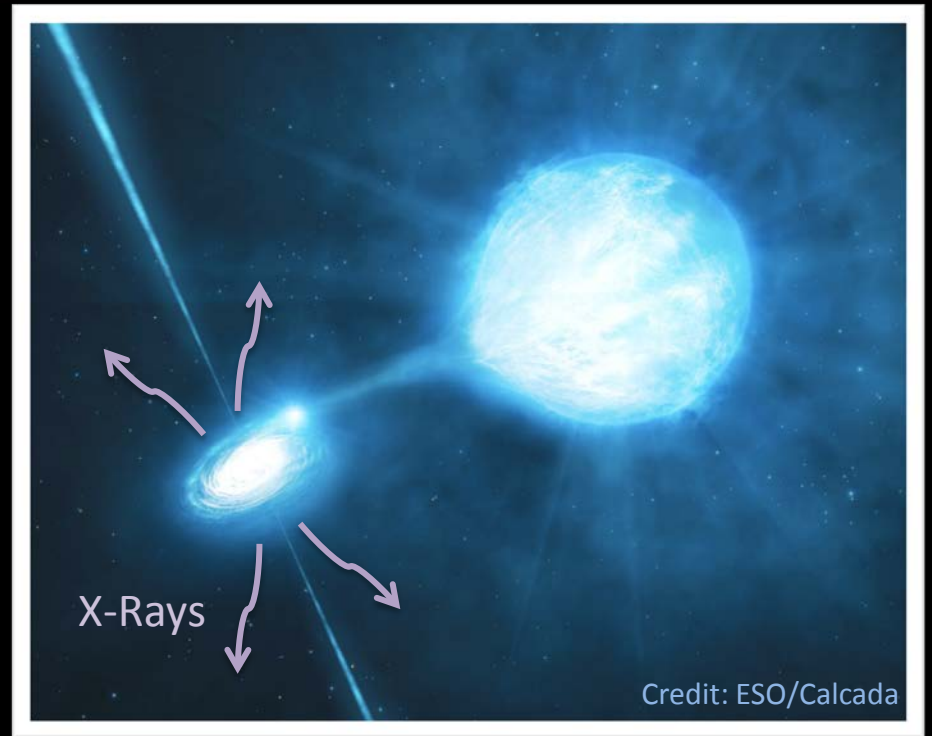
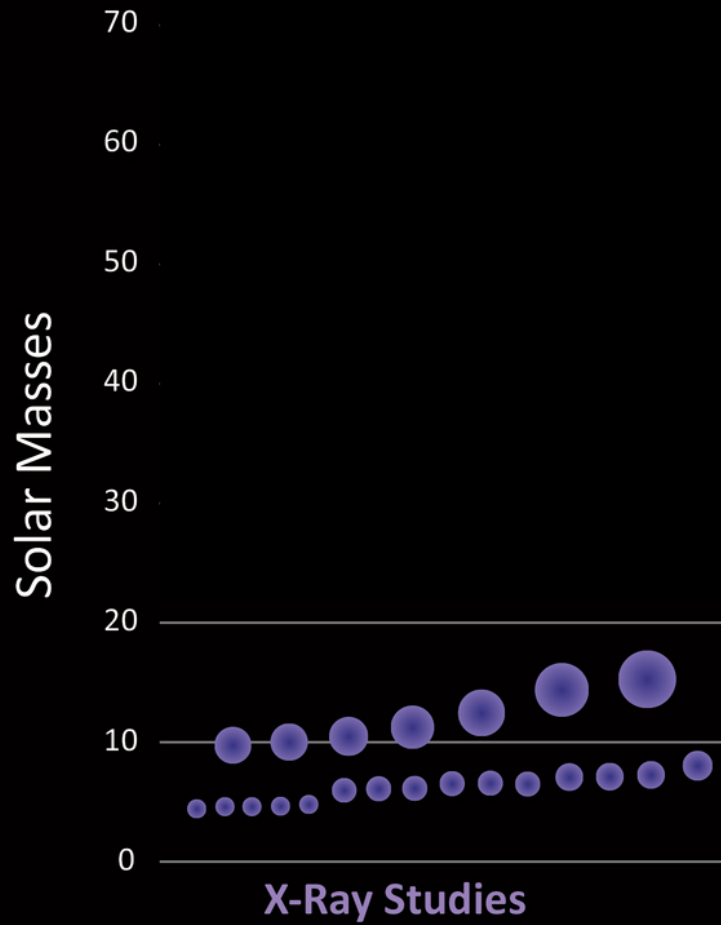


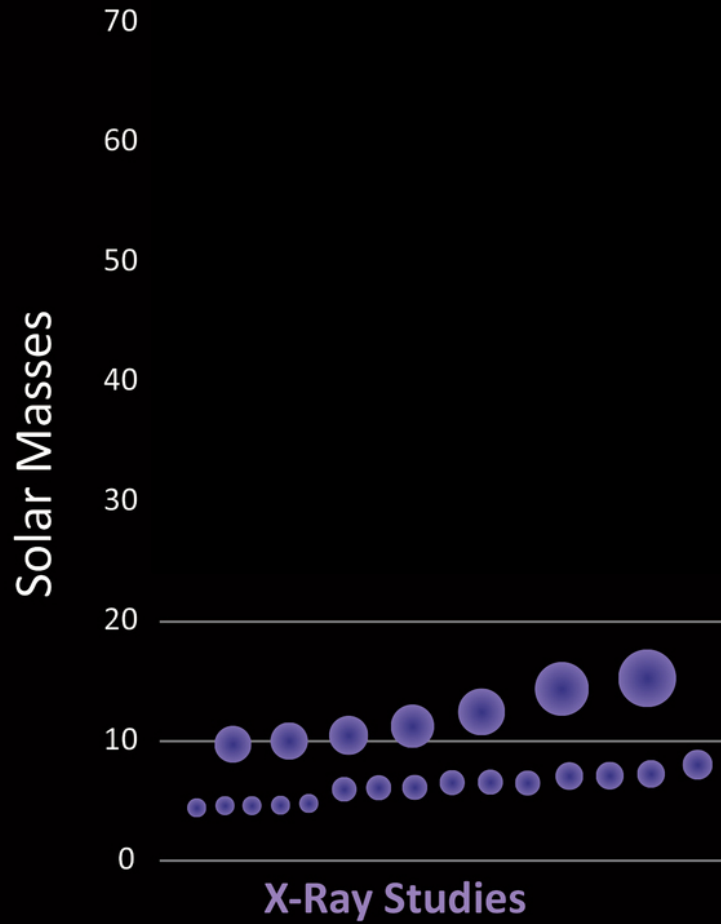
Background design E. Buunk. Credit: NASA/Paresce



**Selma E. de Mink**  
**Black-hole Binary Formation\***  
**Sackler Conference 2018**

\*Field, Non  
Common Envelope





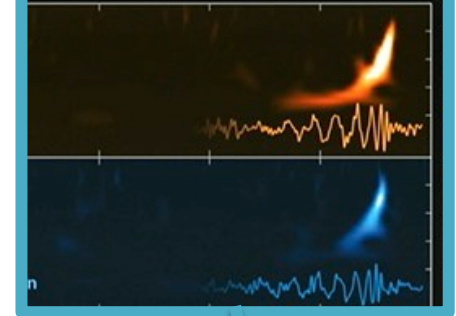
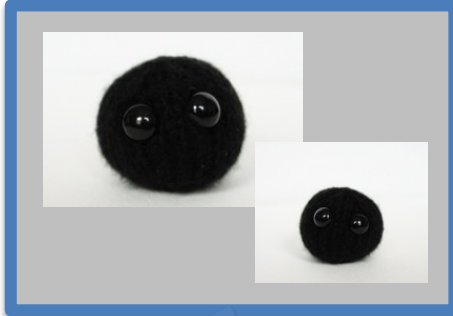
# Intuition for the timescales

4. Birth of progenitors stars

3. Formation binary black holes

2. Merger

1. Detection of signal



*Time*

GW detections still probe relatively local events for now, but the already probed stellar endpoints at very substantial redshifts.

Sept 14  
2015

# Probing how stars die across cosmic time

(A)  
Stellar Origin

(B)  
Primordial

*dark  
ages*

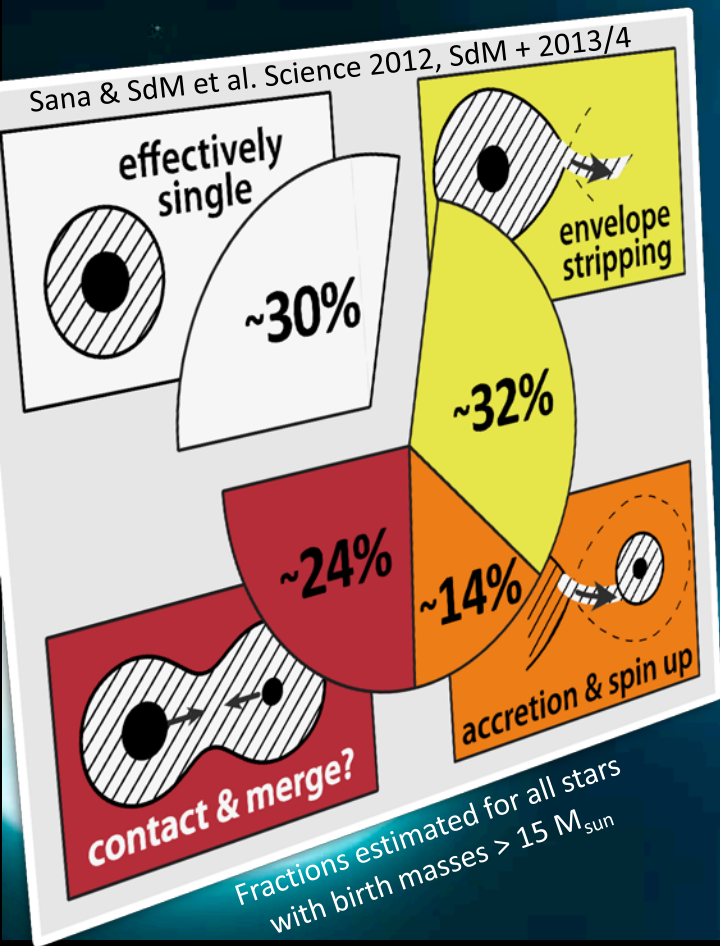


You are  
here

→  
"Cosmic time"  
(not to scale)

e.g. Carr & Hawking (1974), Carr  
(1975, 1976), Garcia-Bellido et al.  
(1996), Khlopov (2010), Frampton et  
al. (2010), Blais et al. (2002), ...

# This is not exotic

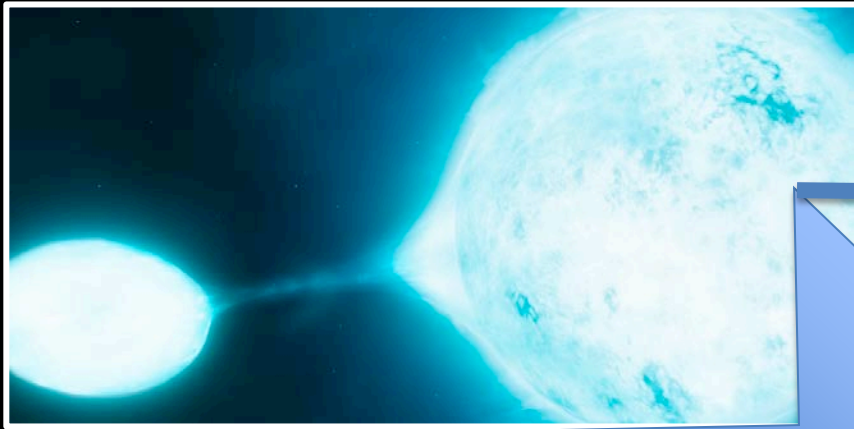


Initial distributions of

- Masses
- Mass ratios
- Separations/Periods
- ...
- Cosmic Starformation history  $f(Z, z)$

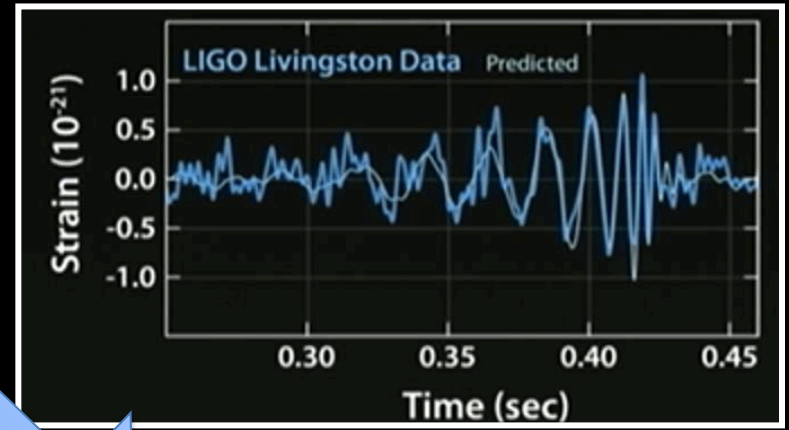
cf. Abt+78, Kobulnicky+Fryer07, Mason+09,  
Chini+12, Sana,SdM+12, Sana,deKoter,SdM+13, Kobulnicky+14,  
Dunstall+w/SdM15, Moe+16, Almeida+w/SdM17, ...

*This is not Exotic ...*



?

*... but this is.*



Other  
exciting  
objects



# Formation Channels



# Why is it so hard to make close double compact objects



1. *“Separation Challenge”*

2. *“Mass Challenge” (for BHs)*

# Why is it so hard to make close double compact objects

## 1. "Separation Challenge"

*How to get Black holes close enough to coalesce in a Hubble time?*

$10 R$



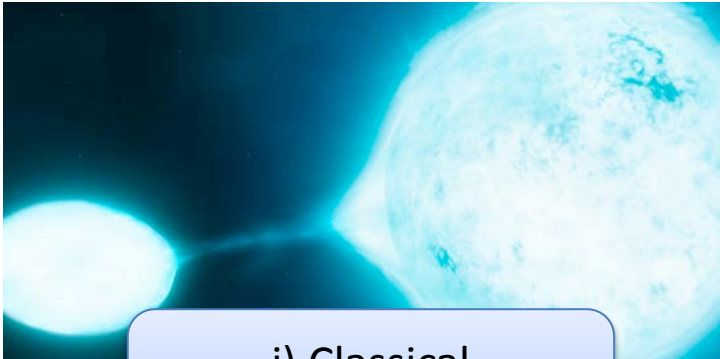
## 2. "Mass Challenge"

*How to avoid excessive Mass loss?*

# Formation Channels

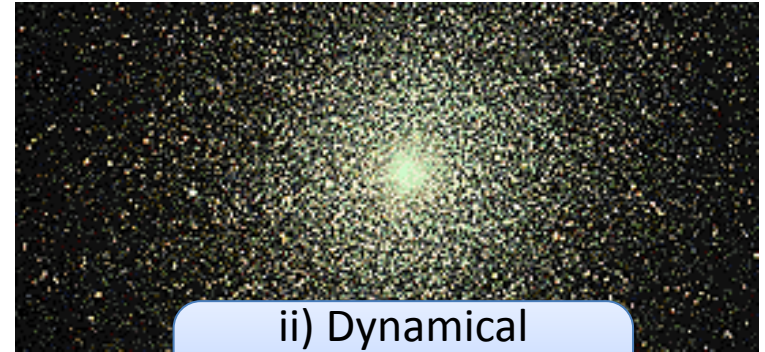
## Traditional Division

### “Evolutionary Channels” (Field)



i) Classical  
(Common Envelope)

### “Dynamical Channels”

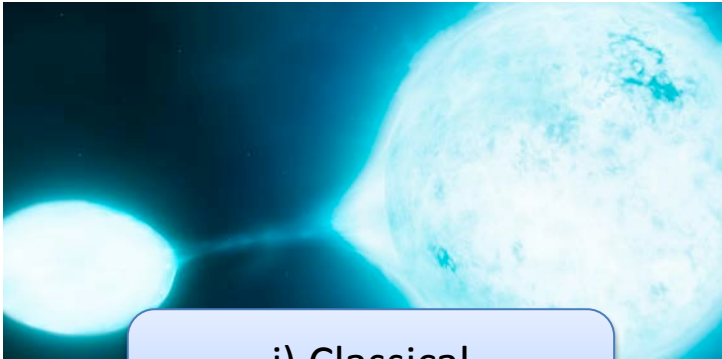


ii) Dynamical  
formation in  
massive star clusters

Cf. talks by Mandel, Rasio, Naoz, Shoemaker, Holz, ...

# Formation Channels

## “Evolutionary Channels” (Field)



i) Classical  
(Common Envelope)

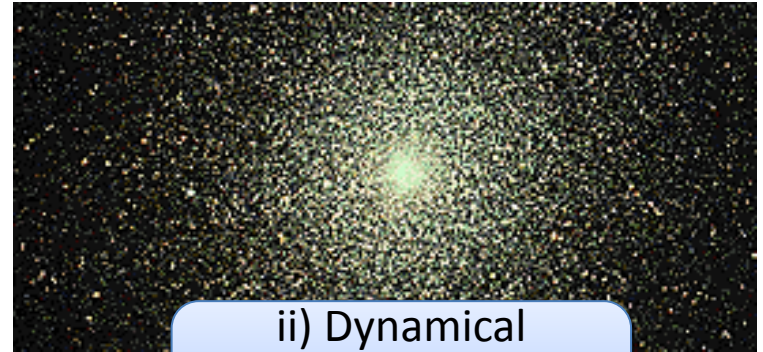
iii) Chemically  
Homogeneous  
Channel

vii) Stable non-  
conservative  
mass transfer

vi) Triples

...?...

## “Dynamical Channels”



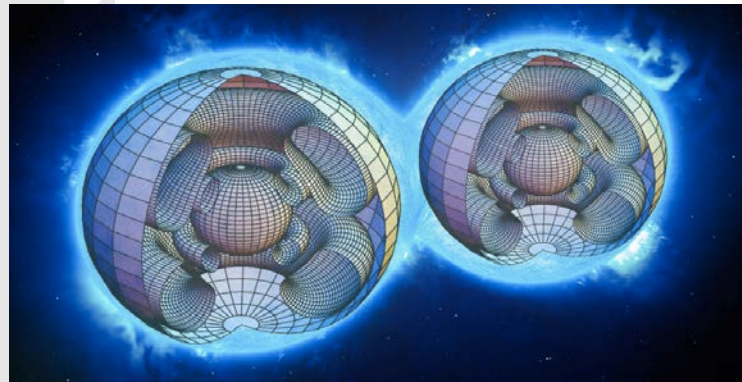
ii) Dynamical  
formation in  
massive star clusters

iv) Kozai  
resonance with  
SMBH

v) In gas disk of  
Active Galactic  
Nuclei

*See talks by Mandel, Rasio, Naoz, Shoemaker, Holz, ...*

# Chemically Homogeneous Channel



“Case M Evolution”,  
“The Rotational Channel”,  
“Tidally Induced Mixing Channel”

de Mink et al. (2008, 2009), Mandel & de Mink (2016), Song et al. 2016;  
Marchant et al. (2016), de Mink & Mandel (2016), Marchant et al. (2017)  
Marchant et al. (to be subm.)

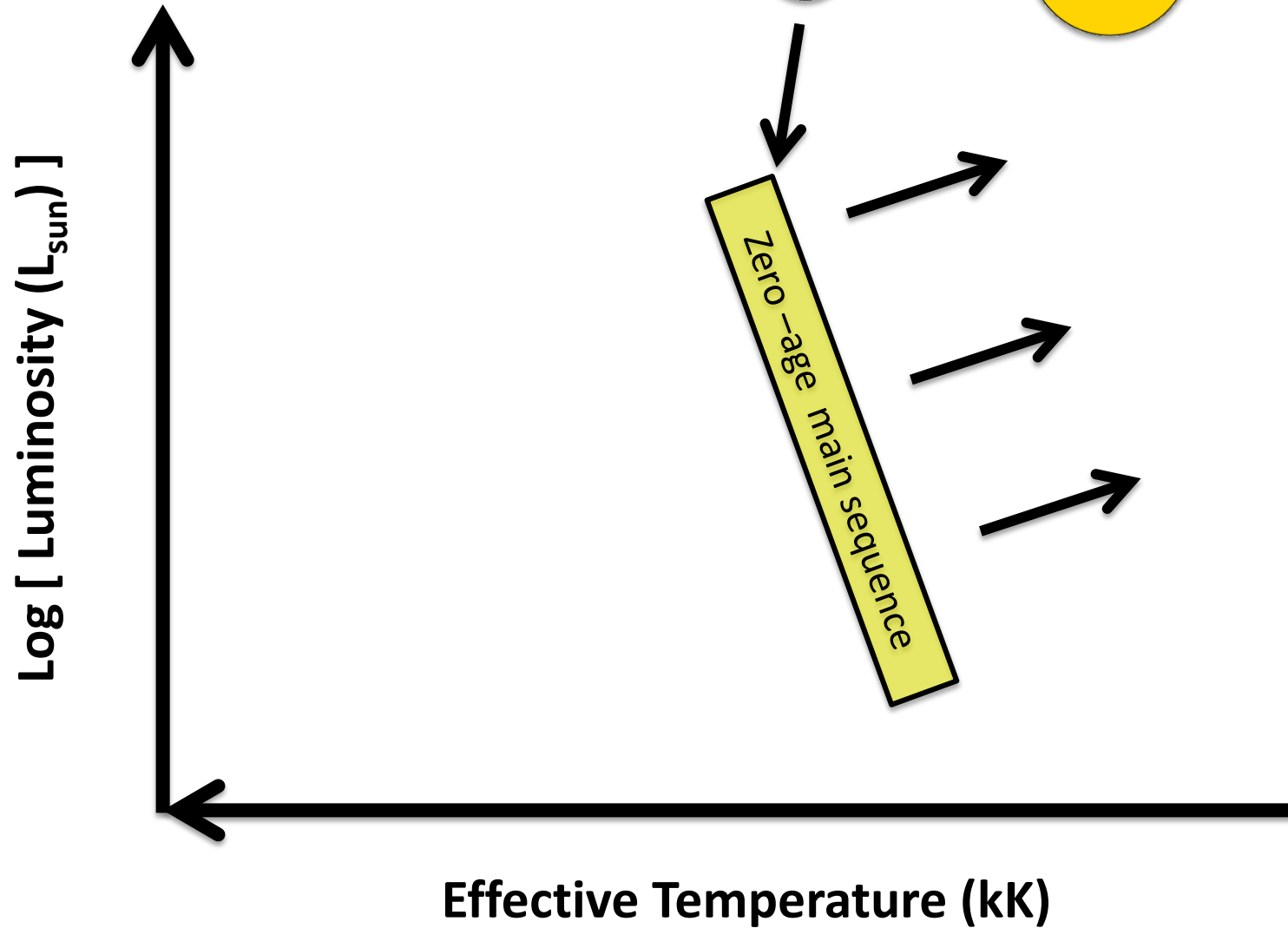
Primary Collaborators:

Ilya Mandel (2016), Matteo Cantiello (2009)

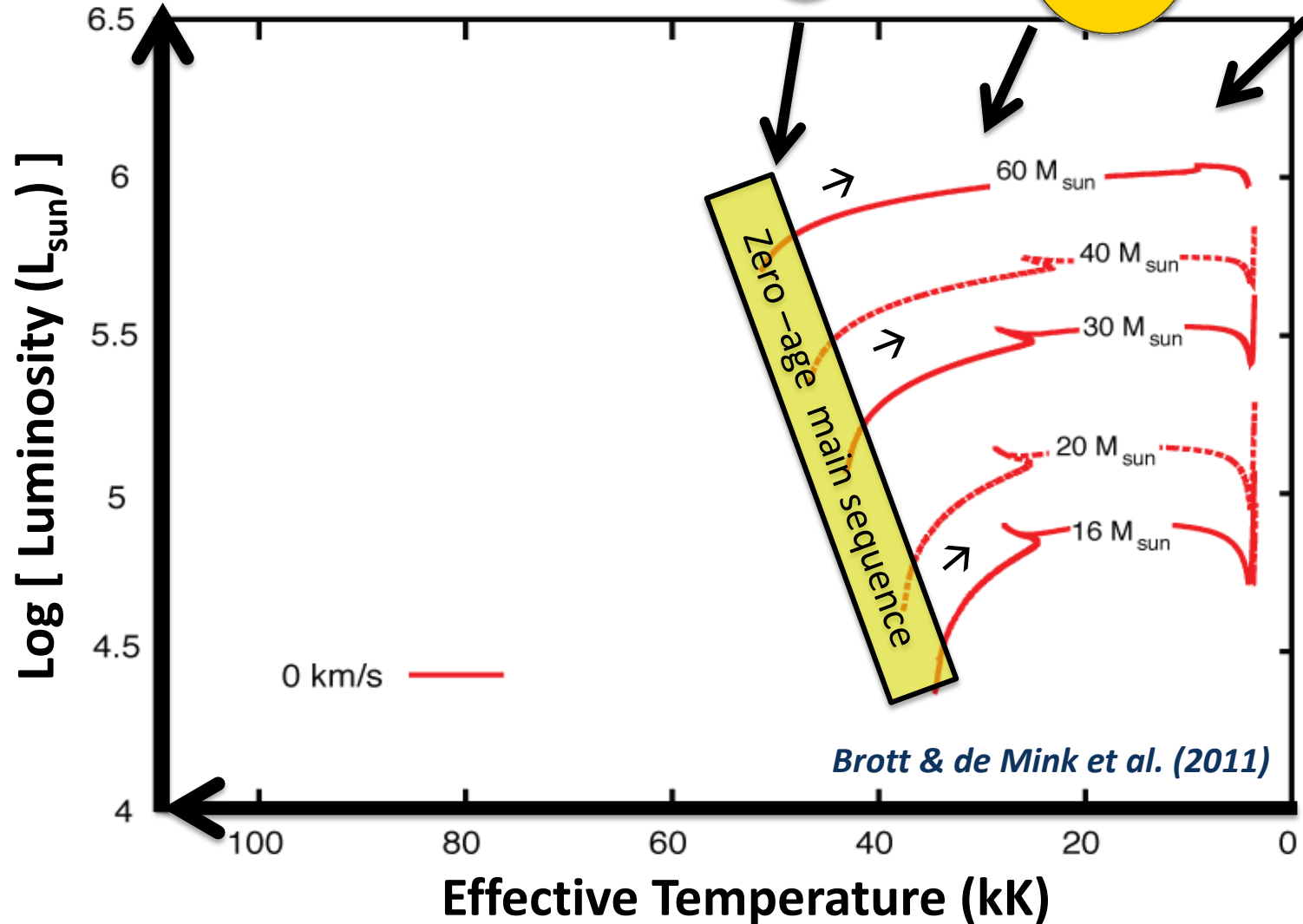


**Pablo  
Marchant**

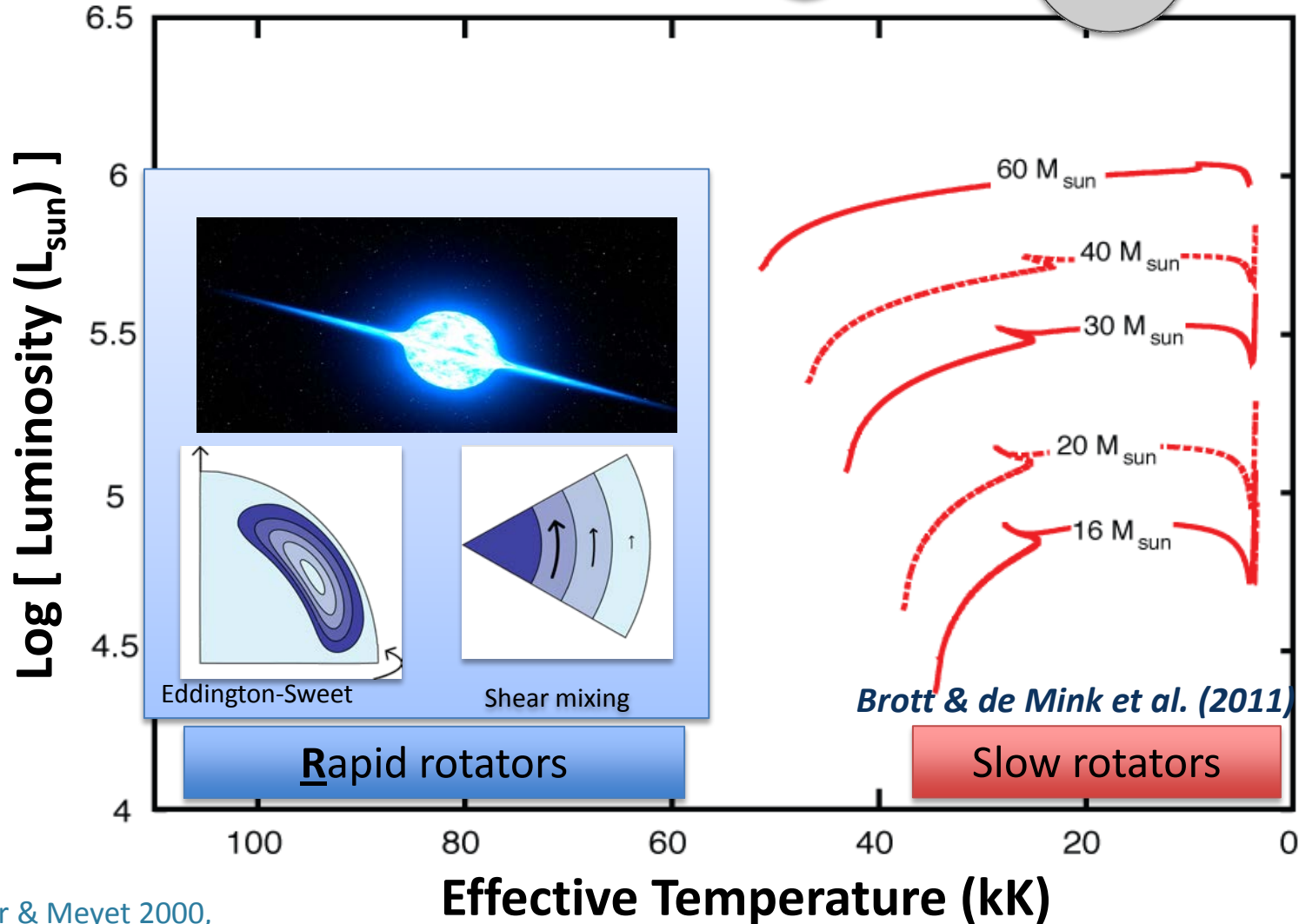
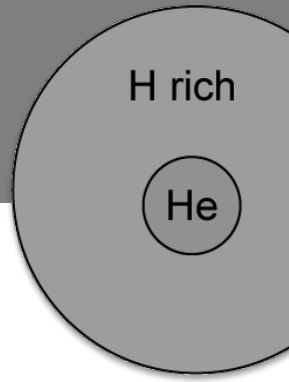
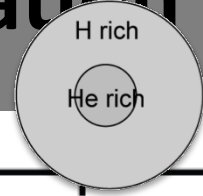
# Single stars with Rotation



# Single stars with Rotation



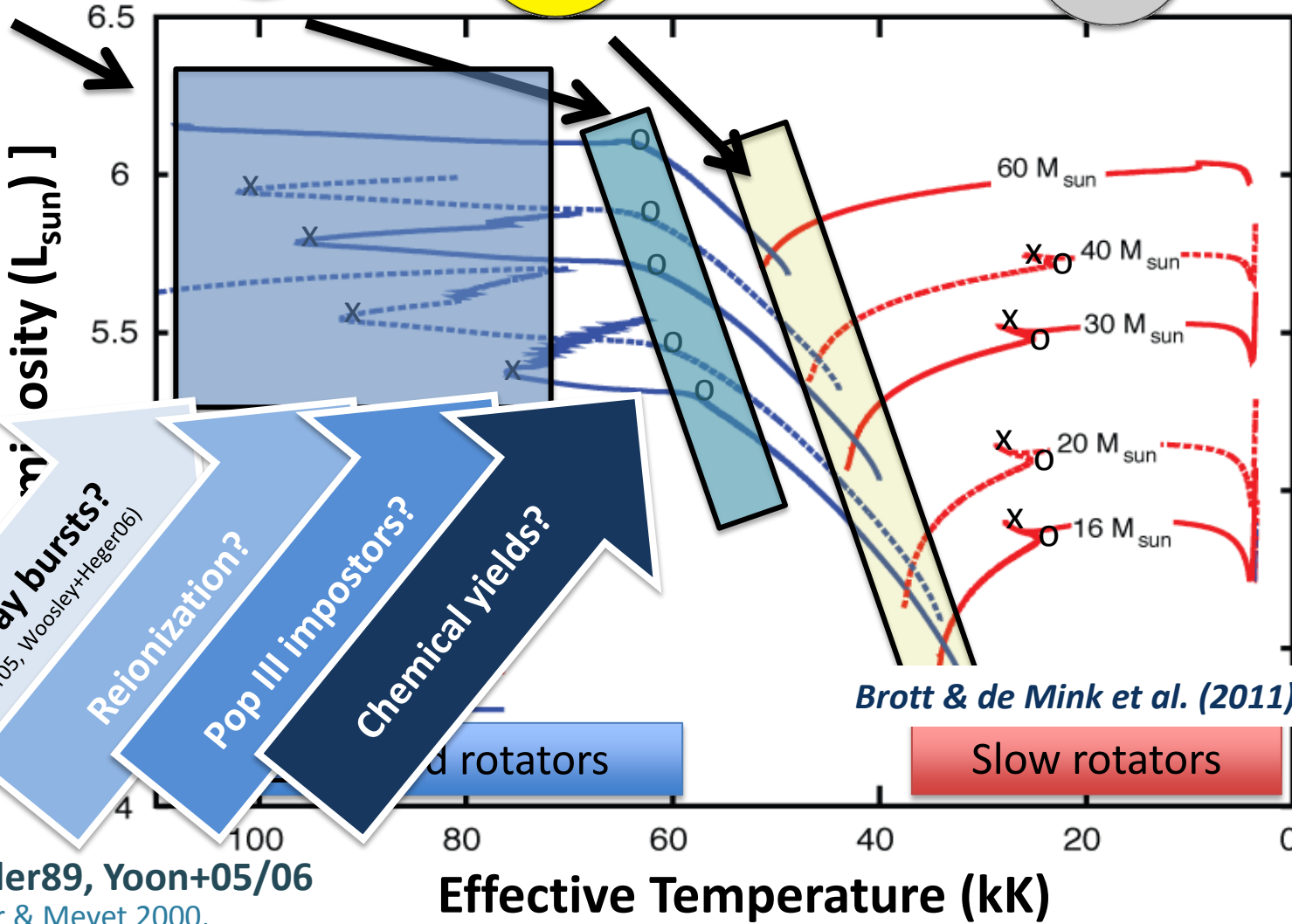
# Single stars with Rotation



Maeder & Meyet 2000,  
 Heger et al 2000, Yoon et al 2006, Hirschi 2007, Ekstroem et al 2012, Georgy et al. 2014, Kohler et al. 2015

$Z = 0.004$

# Single stars with Rotation



**Gamma-ray bursts?**  
(Yoon+Langer05, Woosley+Heger06)

**Reionization?**

**Pop III impostors?**

**Chemical yields?**

*Brott & de Mink et al. (2011)*

**Maeder89, Yoon+05/06**  
Maeder & Meyet 2000,

Heger et al 2000, Yoon et al 2006, Hirschi 2007, Ekstroem et al 2012, Georgy et al. 2014, Kohler et al. 2015

$Z = 0.004$

# What if you could do this in a close binary?

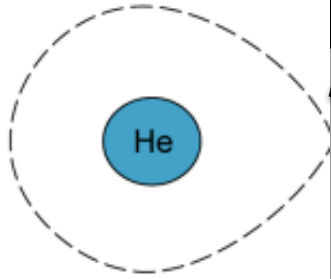
Standard Evolution

Chemically Homogeneous

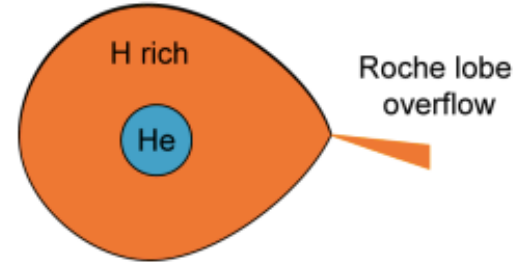
# Full Evolutionary Calculations

De Mink + 2009

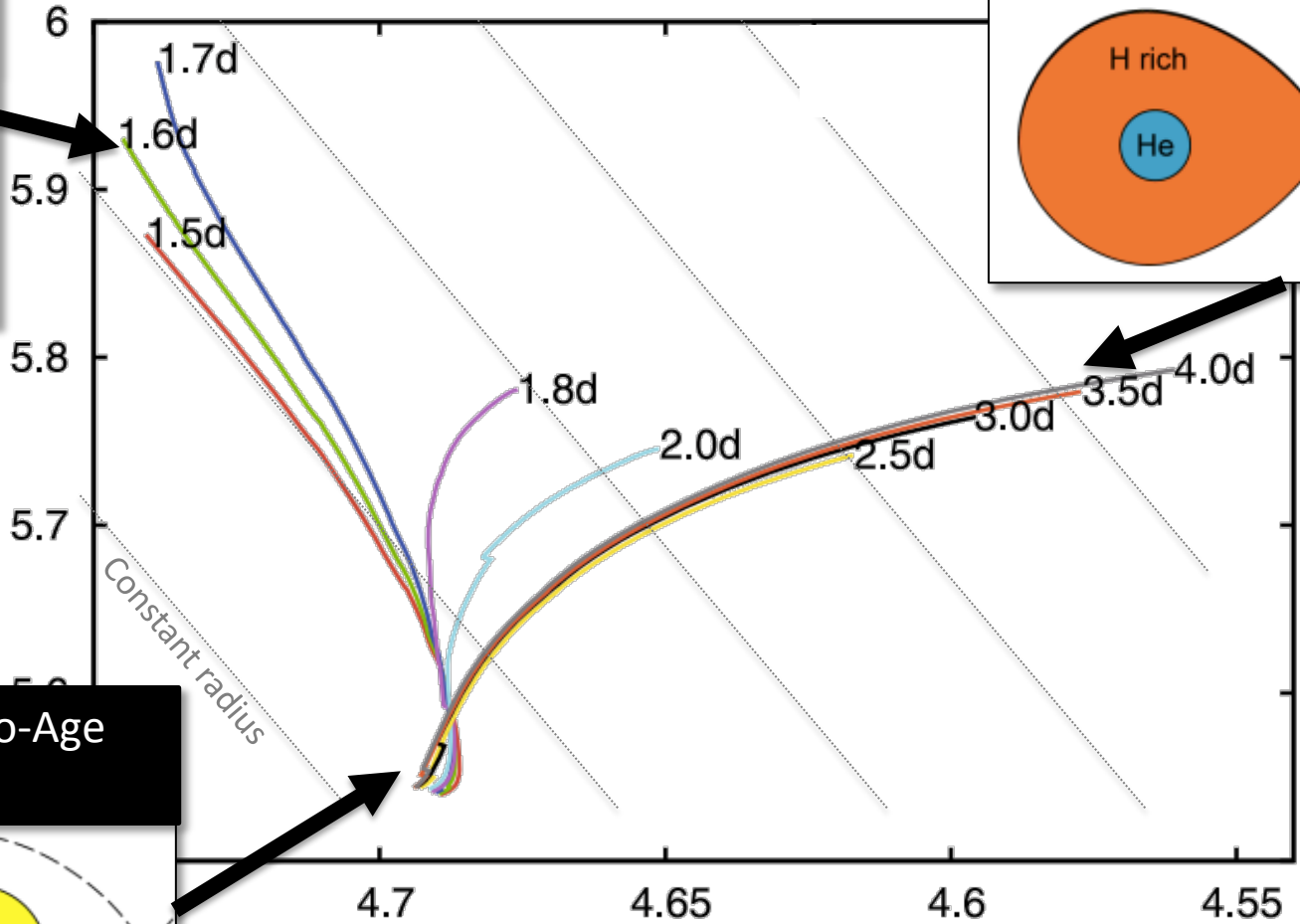
2b)  
Homogenous



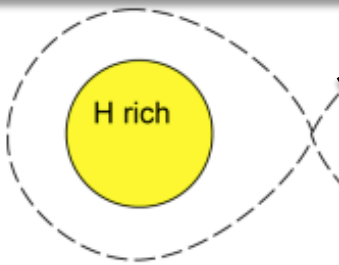
2a) Normal  
Evolution



g Luminosity



1. Zero-Age

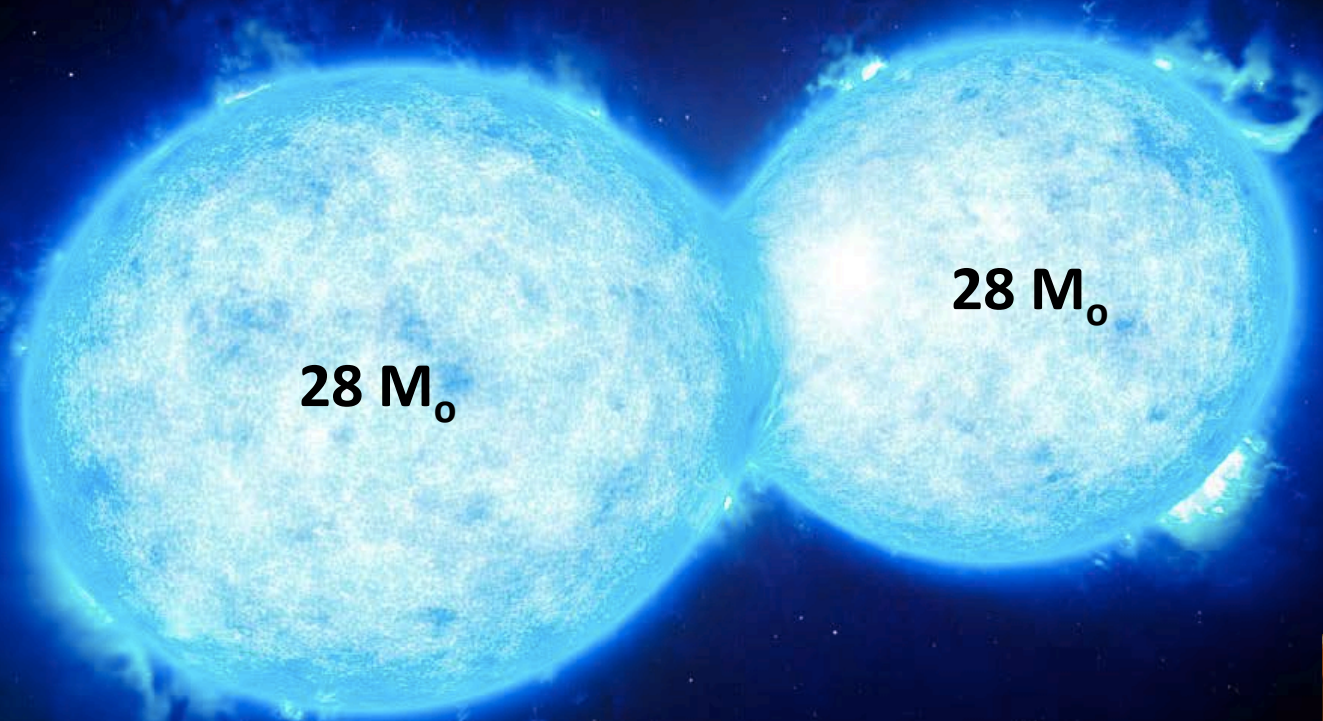


Log Effective Temperature (kK)

# Do such tight binaries exist?

**VFTS 352 - Orbit: 1.1 Day**

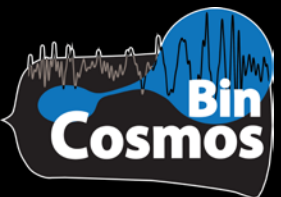
**A case for (partial)Chemically Homogeneous Evolution?**



Almeida, Sana, de Mink et al. (2015)  
Image Credit:ESO/L. Calçada



Leonardo  
Almeida

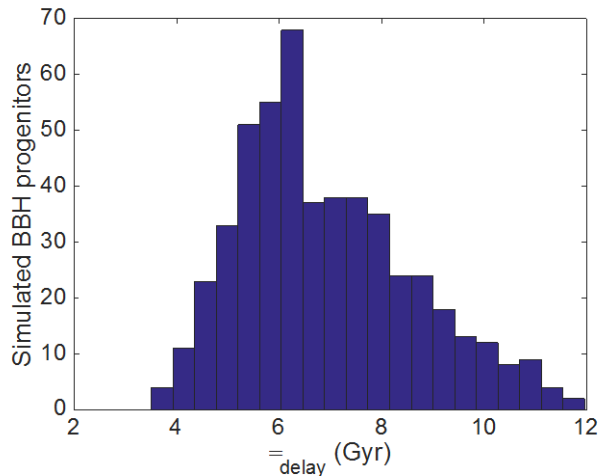


# Predictions for the Chemically Homogeneous Channel

(Take these with a grain of salt)

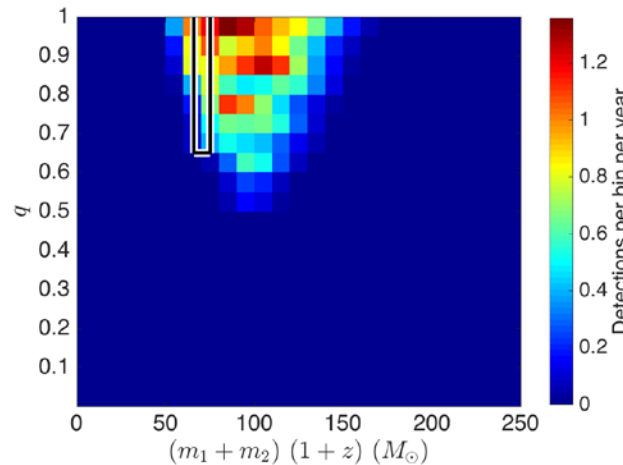
## Delay Time

Mandel & de Mink (2016)



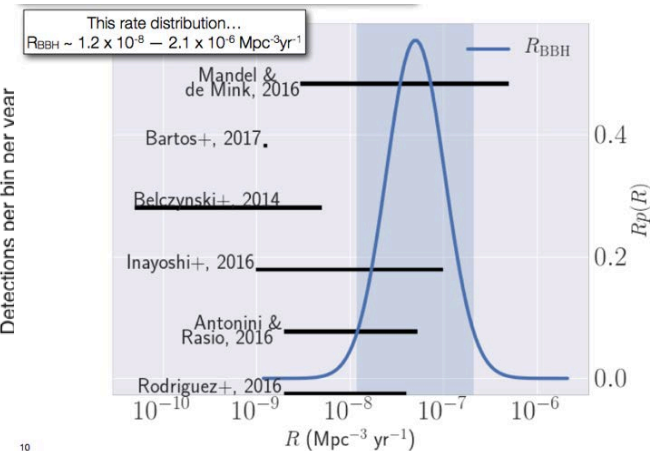
## Masses

de Mink & Mandel (2016)



## Rates

DCC: LIGO-G1800370



## Open Questions:

- **What about Spins?** - *Open question, hard to predict*
- **Lower mass events?** – *Probably only massive*
- **What about side products?** - ....

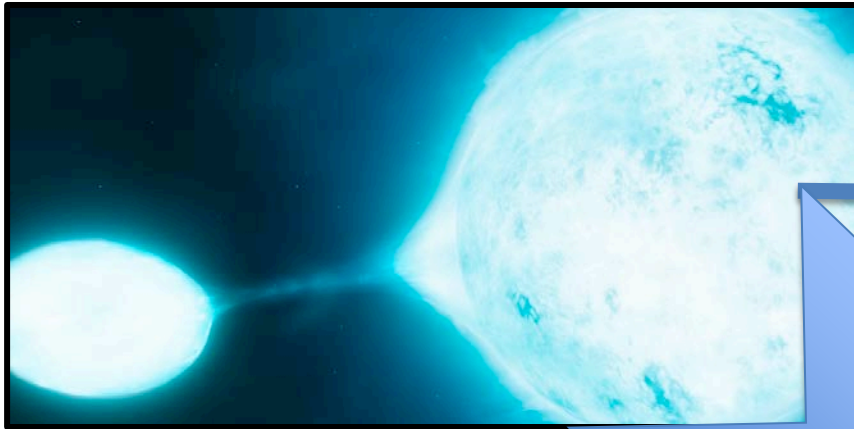
Cf. Marchant et al. (2016)



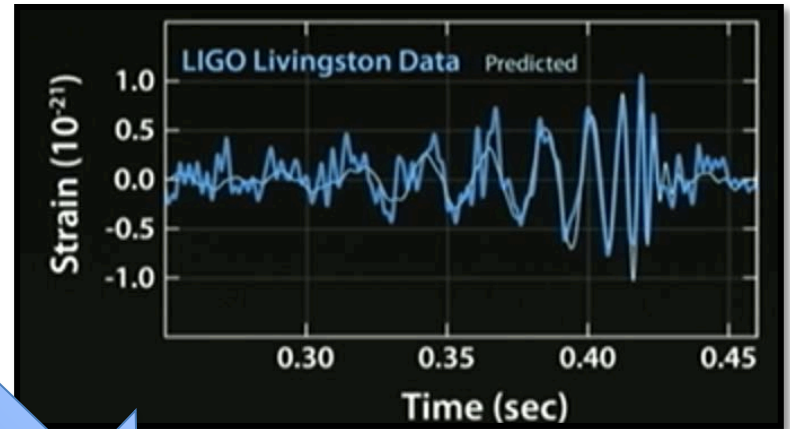
# How to Proceed from here



# How to Proceed from here?



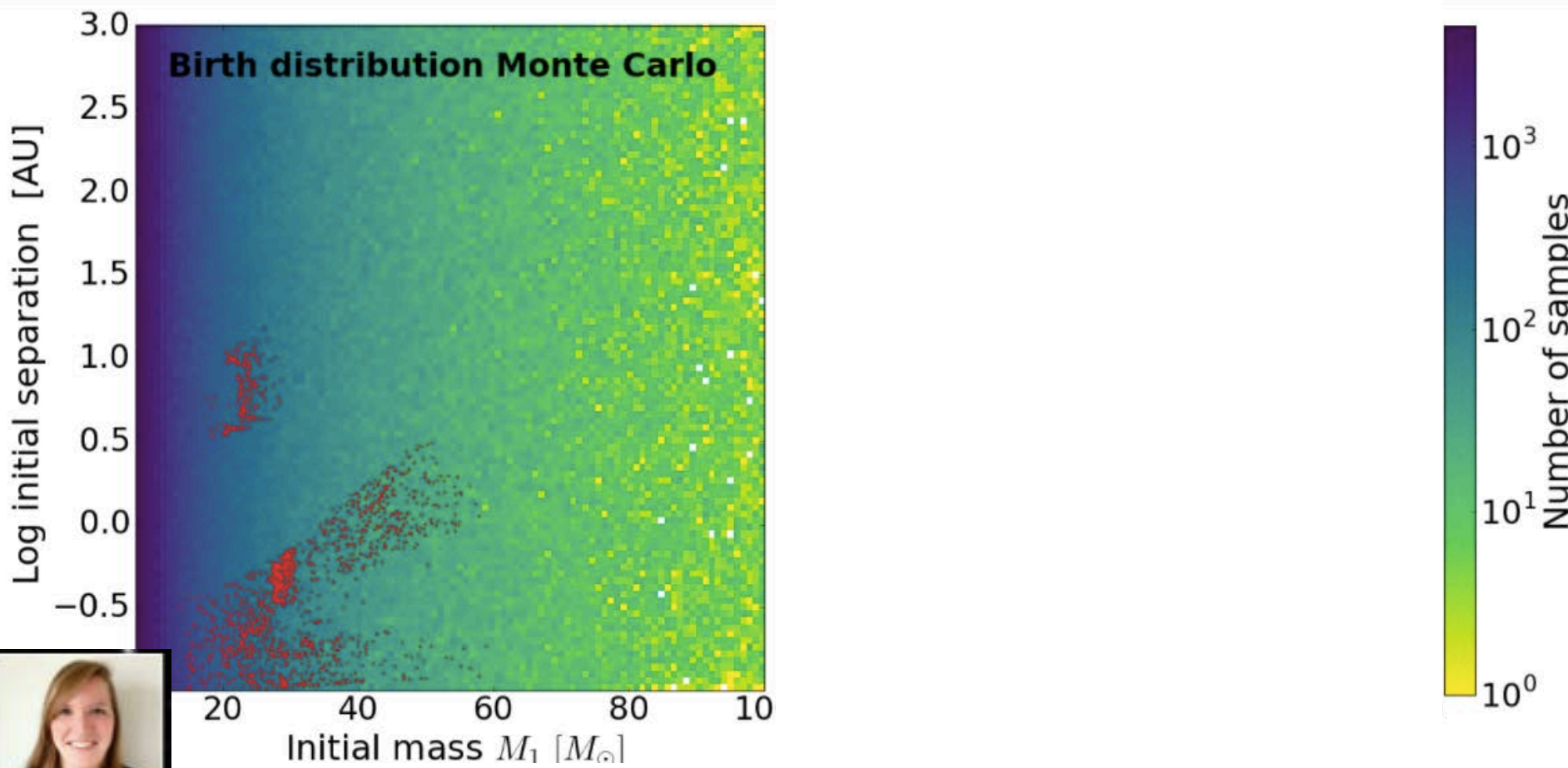
?



Side  
Products

- *Runaway stars*
- *Overcontact binaries*
- *Stellar mergers*
- *(Ultra Luminous) X-ray Sources*
- *(Pair Instability) Supernovae*
- ....

# Challenge of Simulating Rare Events (such as BH-NS systems)



BH-NS systems

1 Million Binaries  $\sim$  90 CPU hours, w/ Compas

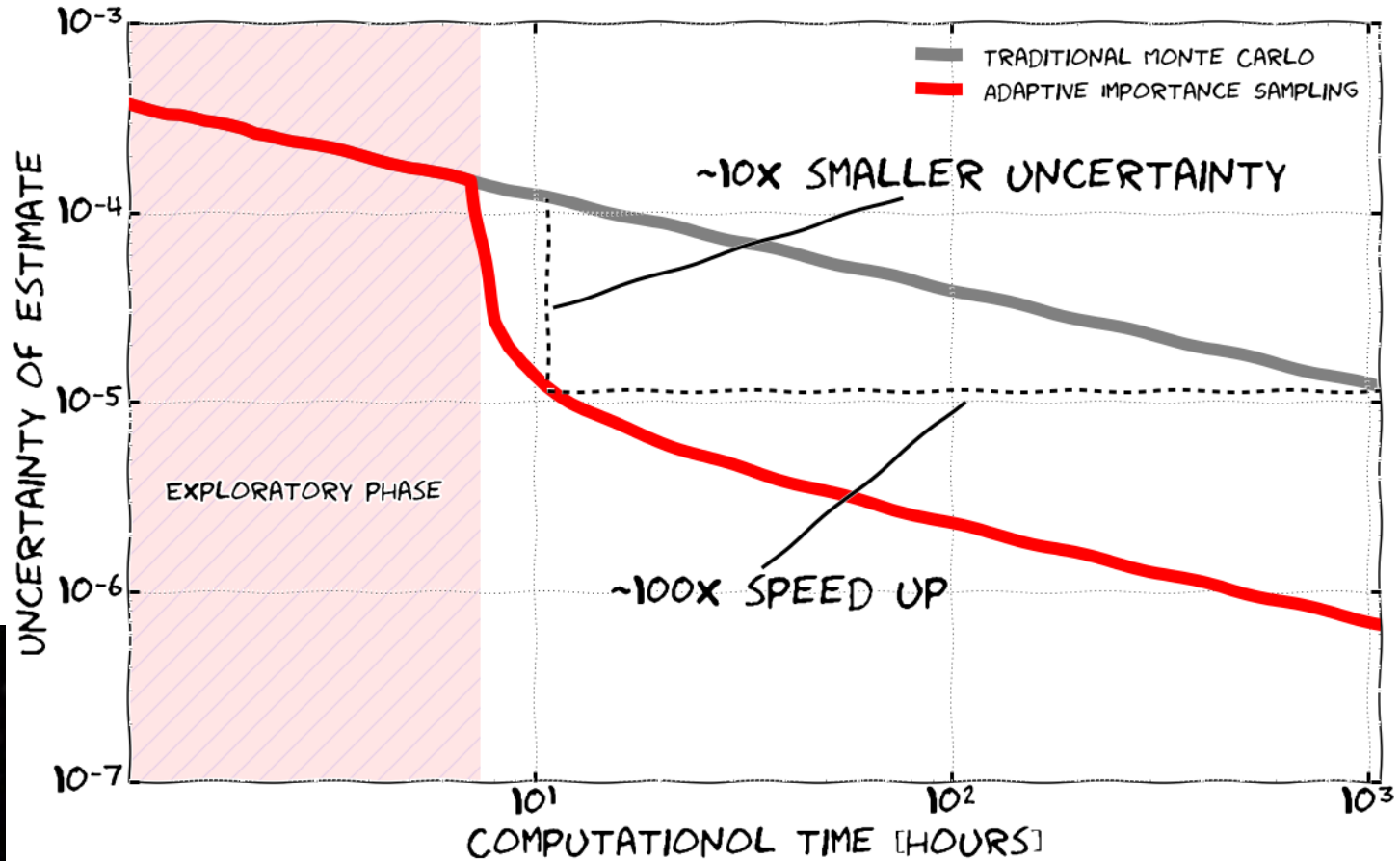
w/ Gair, Justham, Baret, Justham, Mandel, Stevenson, Vigna-Gomes



Floor Broekgaarden

# Challenge of Simulating Rare Events

Broekgaarden et al. in prep.



Floor Broekgaarden

# How not to make a binary black hole

# Side Products (1): Can Gaia Help?

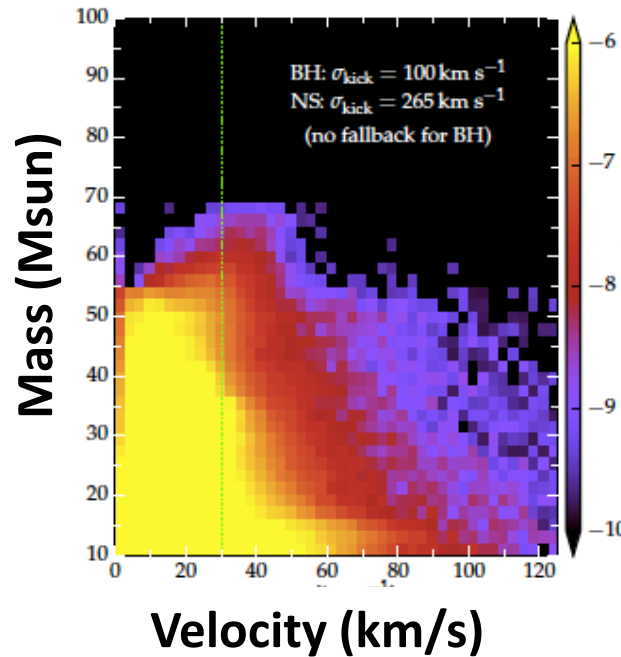
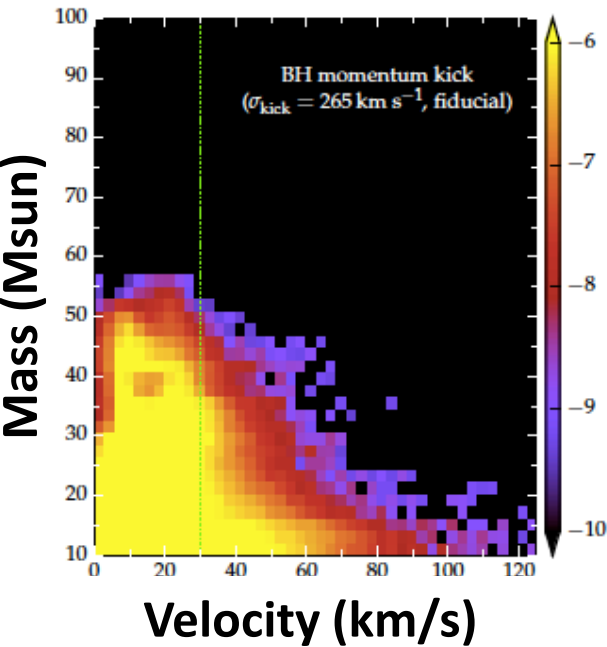


Mathieu  
Renzo

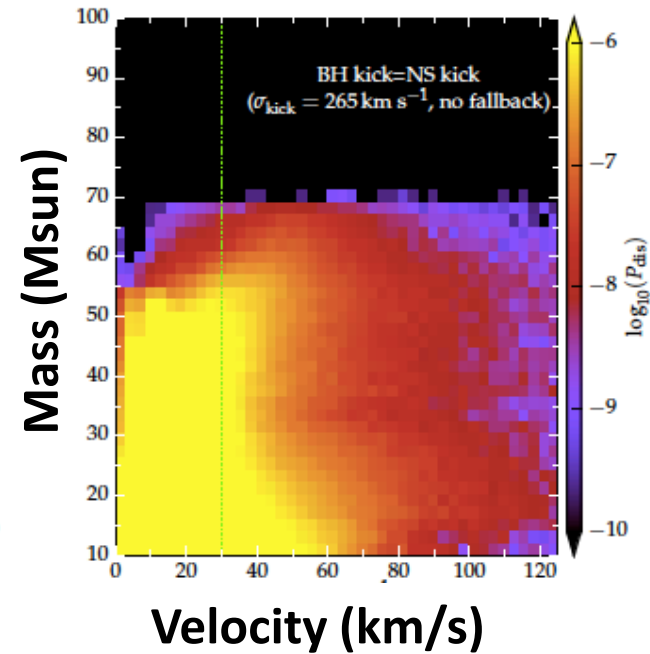
## Runaway and Walkaway Stars

Renzo et al. w/SdM 2018

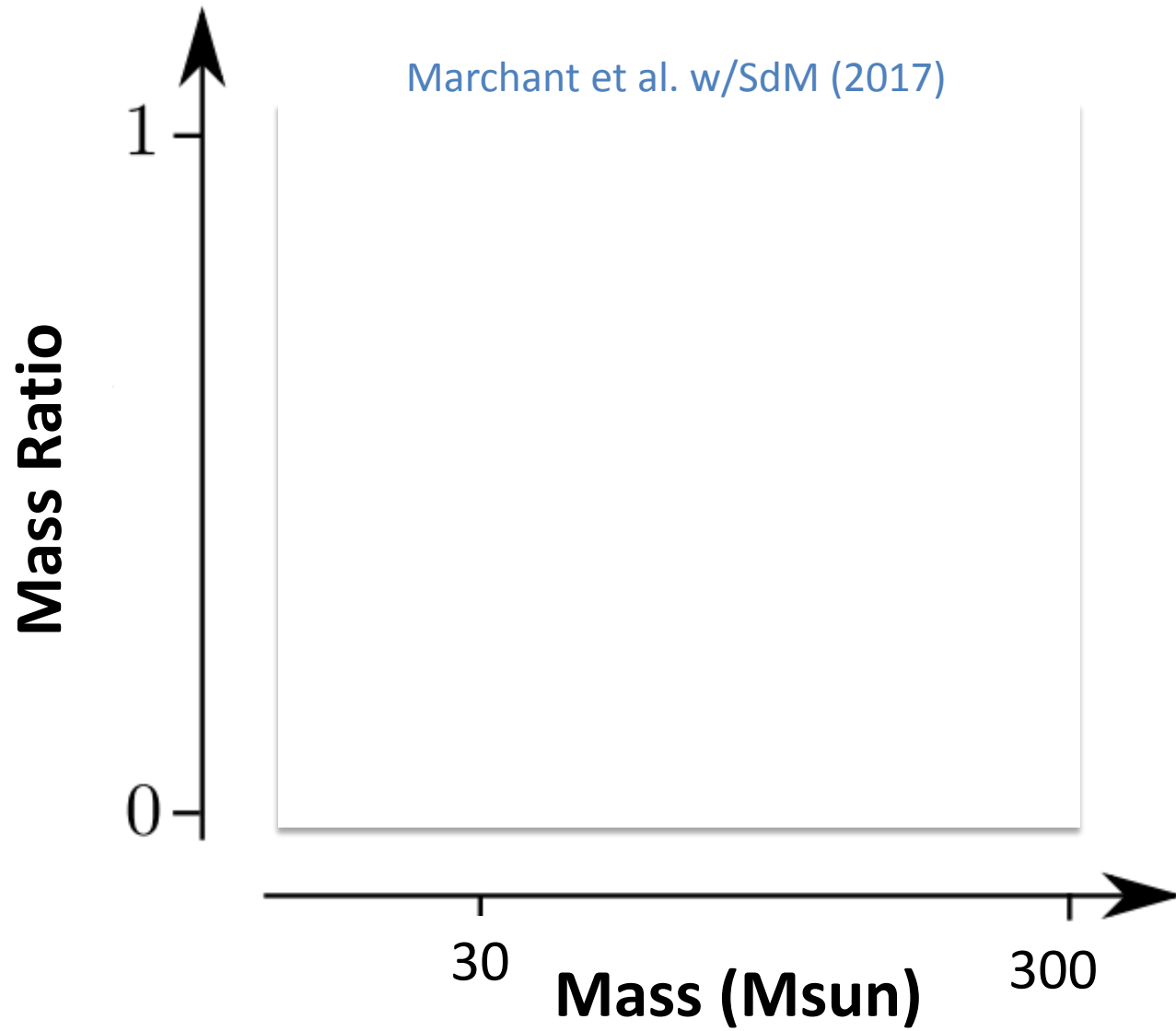
*“Small BH kicks”*



*“Big BH kicks”*

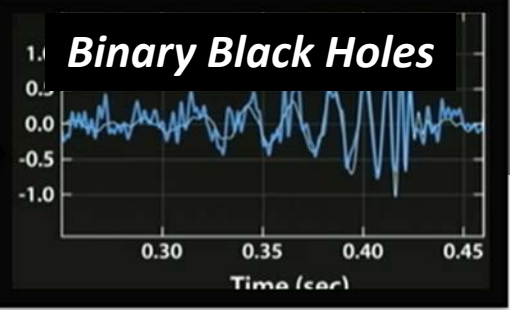


# Side Products (2)



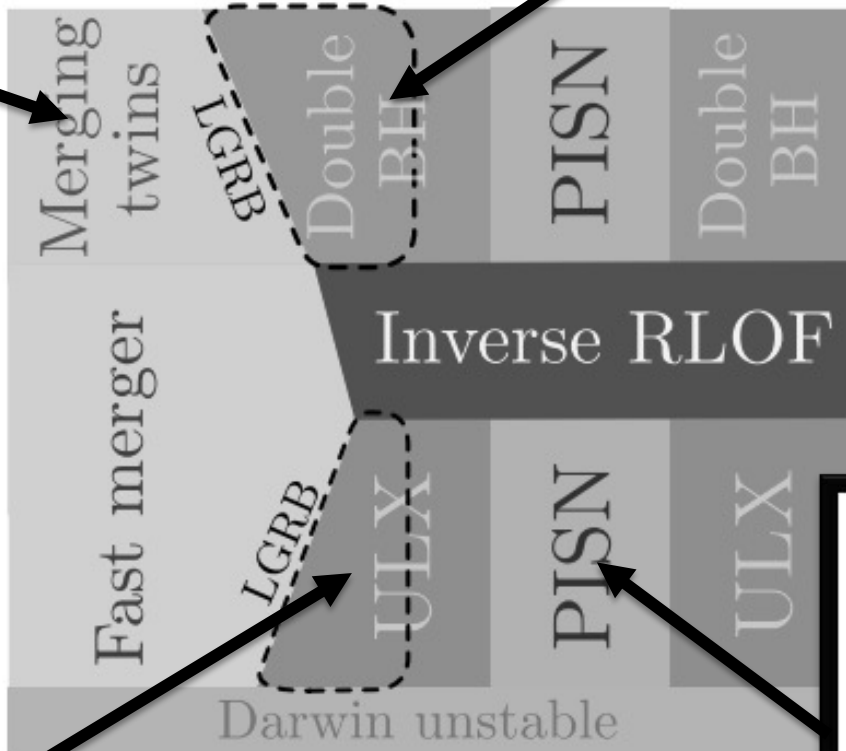
# Side Products

## Binary Black Holes



Marchant et al. w/SdM (2017)

Mass Ratio



Merging twins

LGRB

Double BH

PISN

Double BH

Fast merger

Inverse RLOF

LGRB

ULX

PISN

ULX

Darwin unstable

30

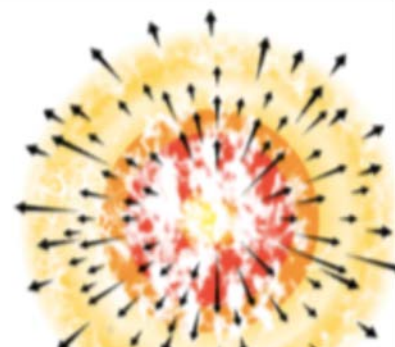
Mass

3

Massive Contact binaries



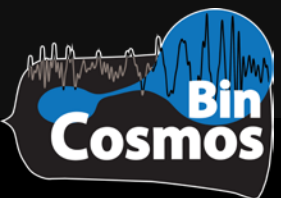
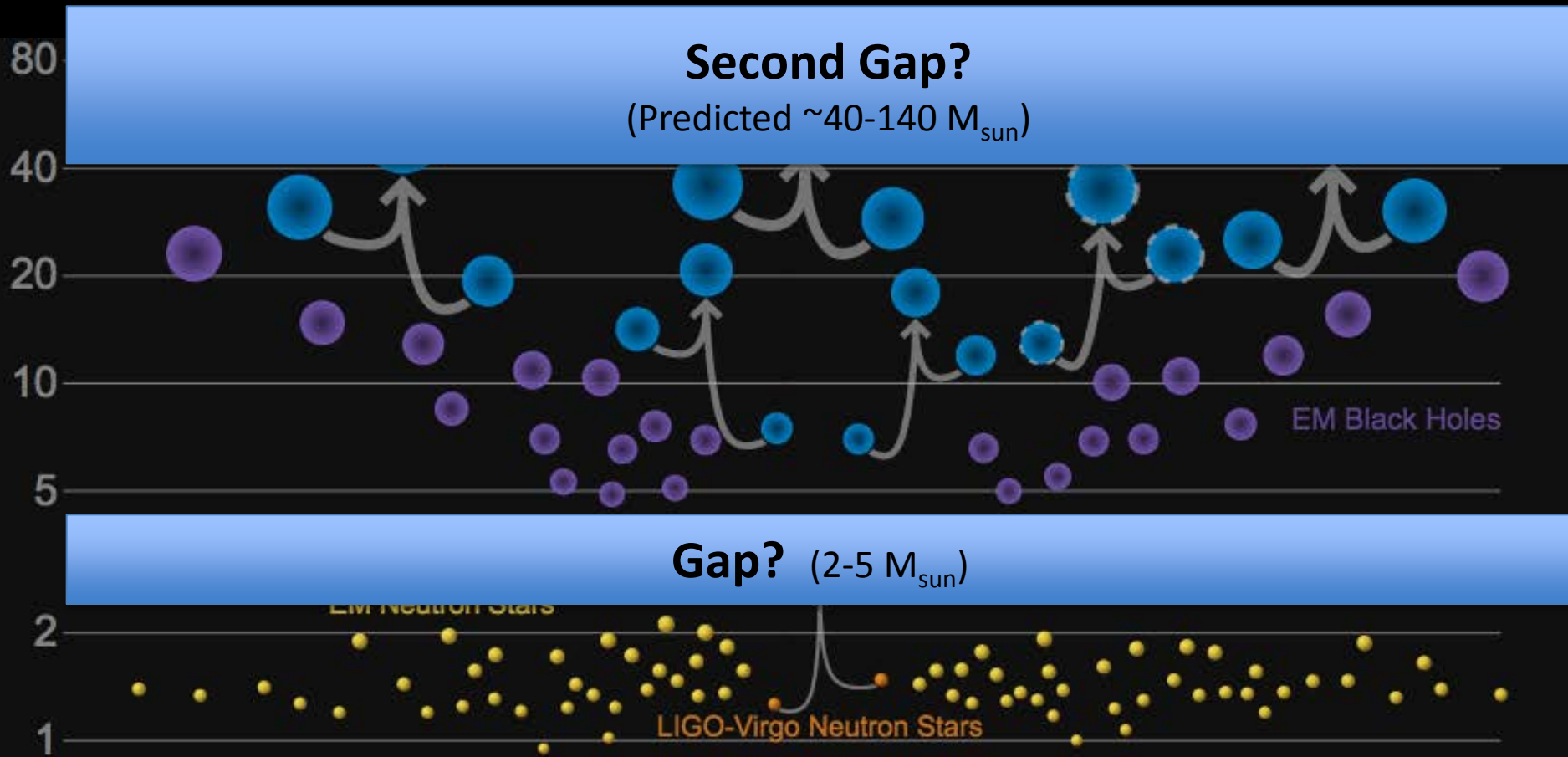
Ultra Luminous X-ray sources

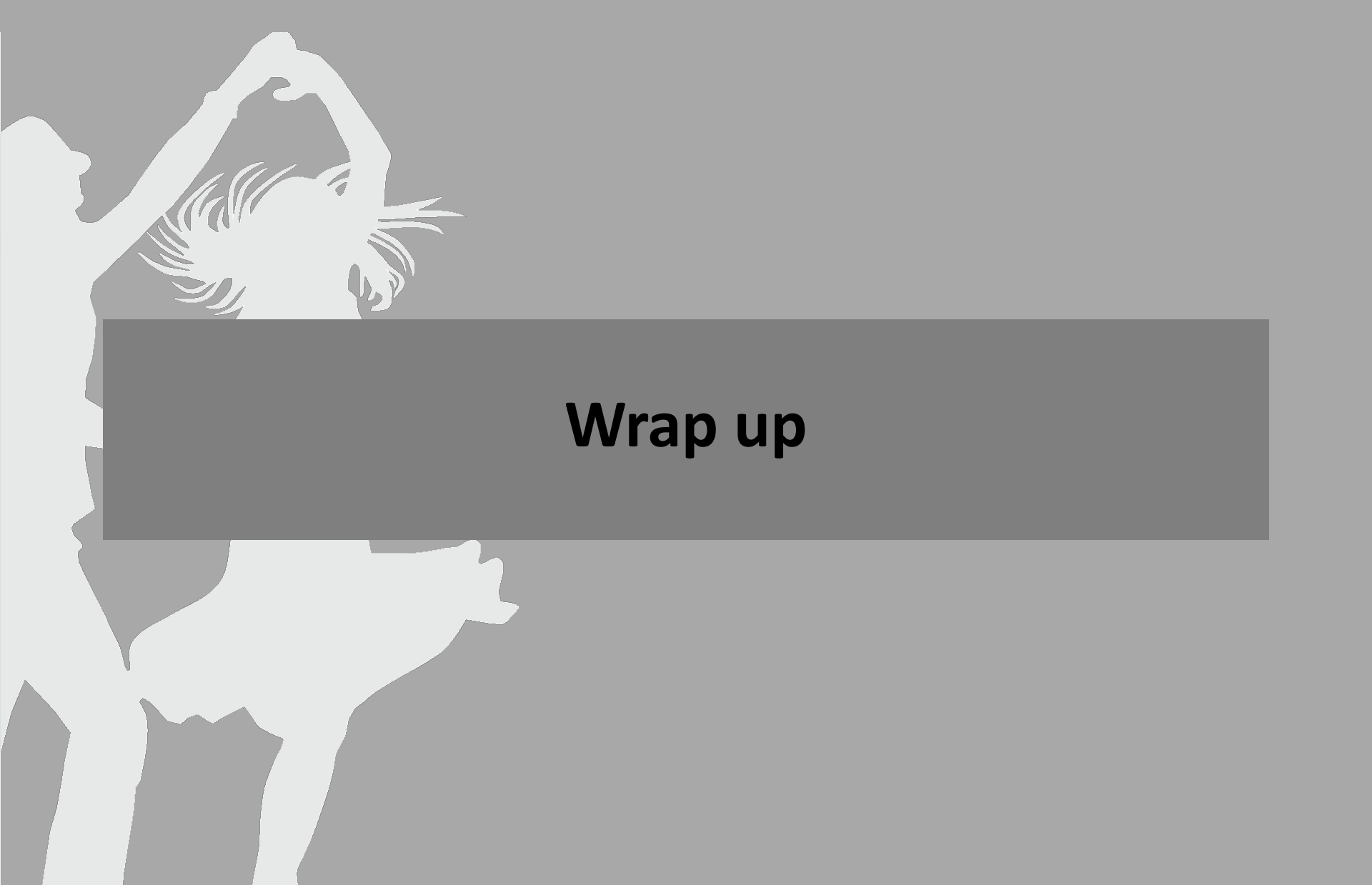


Pair Instability Supernova



Rob Farmer

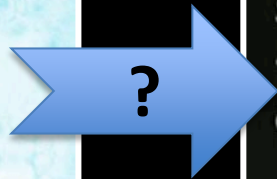
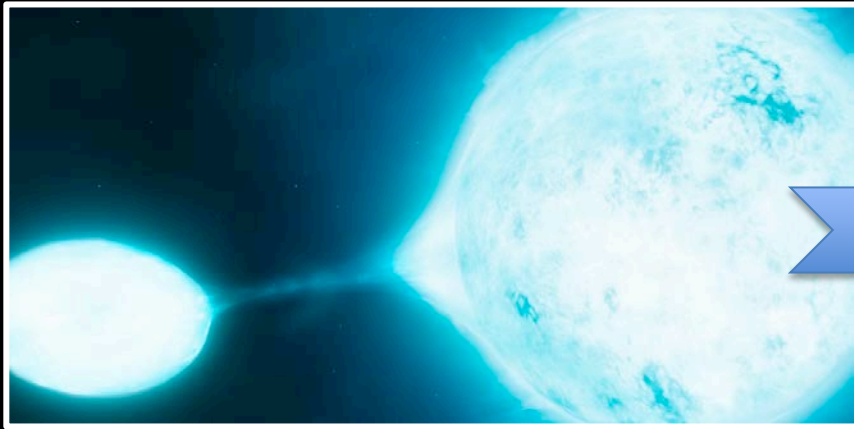




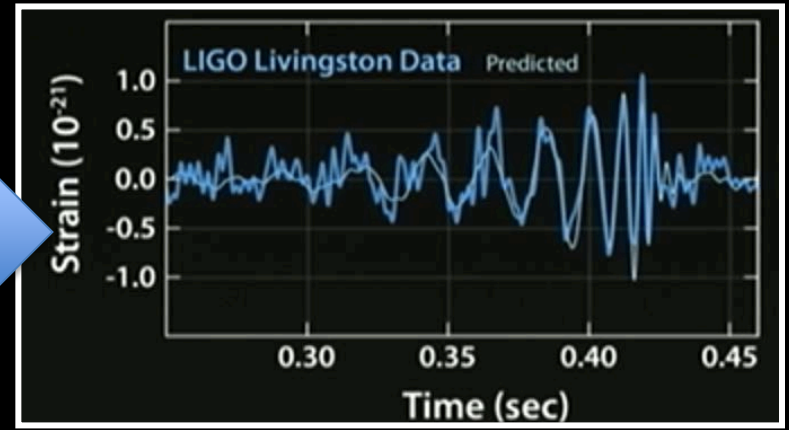
# Wrap up

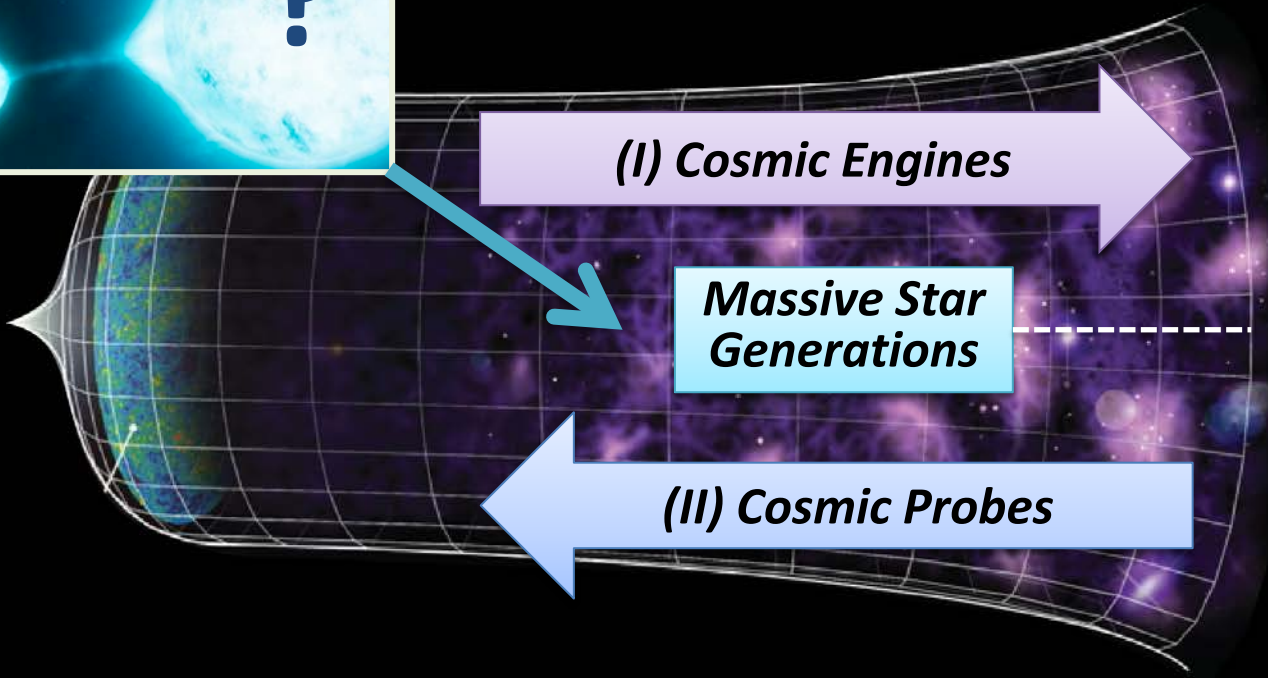


*This is not Exotic ...*



*... but this is.*

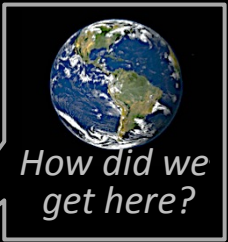




*(I) Cosmic Engines*

*Massive Star  
Generations*

*(II) Cosmic Probes*



How did we  
get here?



Selma de Mink  
PI



Stephen Justham  
Associate '18-



Silvia Toonen  
VENI fellow '17-



Rob Farmer  
Postdoc '17-



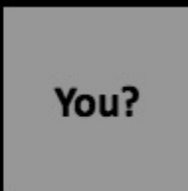
Athira Menon  
Postdoc '18 -



Ylva Götberg  
PhD '18



Manos Zapartas  
PhD '18



You?

...  
...  
PhD '22



Mathieu Renzo  
PhD '19



Eva Laplace  
PhD '21



Floor Broekgaarden  
MSc



Walter van Rossem  
MSc



David Hendriks  
MSc



Karel Temmink  
MSc



Thomas Dodds  
BSc

